Teaching Quantum Mechanics with MATLAB

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Abstract

Among the ideas to be conveyed to students in an introductory quantum mechanics course, we have the pivotal idea championed by Dirac that functions correspond to column vectors (kets $|\alpha\rangle$) and that differential operators correspond to matrices (ket-bras $|\alpha\rangle\langle\beta|$) acting on those vectors. The MATLAB (matrix-laboratory) programming environment is especially useful in conveying these concepts to students because it is geared towards the type of matrix manipulations useful in solving introductory quantum physics problems. In this livescrips, MATLAB codes are included which have been developed at WPI, focusing on 1D problems, to be used in conjunction with Griffiths' introductory text.

Introduction

Two key concepts underpinning quantum physics are the Schrodinger equation and the Born probability equation. In 1930 Dirac introduced bra-ket notation for state vectors and operators [1]. This notation emphasized and clarified the role of inner products and linear function spaces in these two equations and is fundamental to our modern understanding of quantum mechanics. The Schrodinger equation tells us how the state Ψ of a particle evolves in time. In bra-ket notation, it reads

$$(1) \quad i\hbar\frac{d}{dt}|\Psi(t)\rangle=\hat{H}|\Psi(t)\rangle,$$

where \widehat{H} is the Hamiltonian operator and $|\Psi\rangle$ is a ket (or column) vector representing the quantum state of the particle. A side note: in this notes Schrodinger picture is adopted, where state vector evolves in time, and operators corresponding to observables (and hence their eigenbasis of vectors) are time independent (considering closed system described by the time-independent Hamiltonian, with no explicit time dependence). There is another, unitary equivalent picture, the Heisenberg picture, where, state vectors are assued time independent, but instead operators (observables) evolve in time. When a measurement of a physical quantity

A is made on a particle initially in the state $|\Psi\rangle$, the Born equation provides a way to calculate the probability (or probability density for continuous observables) $P(A_o)$ that a particular result A_o is obtained from the measurement. In bra-ket notation, it reads [2]

(2)
$$P(A_a) \sim |\langle A_a | \Psi \rangle|^2$$
,

where $|A_o\rangle$ is the state vector corresponding to the particular result A_o having been measured, $\langle A_o|=|A_o\rangle^\dagger$ is the corresponding bra or row vector and $\langle A_o|\Psi\rangle$ is thus the inner product between $|A_o\rangle$ and $|\Psi\rangle$. In the Dirac formalism, the correspondence between the value of the wavefunction at a point \overrightarrow{x} and time t, $\Psi(\overrightarrow{x},t)$ and the ket $|\Psi(t)\rangle$ state vector is set by the relation $\Psi(\overrightarrow{x},t)=\langle \overrightarrow{x}|\Psi(t)\rangle$, where $|\overrightarrow{x}\rangle$ is the state vector corresponding to the particle being located at \overrightarrow{x} and ket vectors $|\overrightarrow{x}\rangle$ form eigenbasis of coordinate operator. Thus we regard $\Psi(\overrightarrow{x},t)$ as a component of a state vector $|\Psi(t)\rangle$, just as we usually [3] regard $a_i=\overrightarrow{e}_i\cdot\overrightarrow{a}$ as a component of \overrightarrow{a} along the direction \overrightarrow{e}_i . Similarly, we think of the Hamiltonian operator as a matrix

$$(3) \hat{H} = \int d^{3} \overrightarrow{x} | \overrightarrow{x} \rangle \left\{ \frac{1}{2m} \left(\frac{\hbar}{i} \frac{\partial}{\partial \overrightarrow{x}} \right)^{2} + U(\overrightarrow{x}) \right\} \langle \overrightarrow{x} |$$

acting on the space of kets. Acting on state $|\Psi(t)\rangle$ with Hamiltonian (3), using Schrodinger Eq. (1) and multiplying both sides with (time independent!) basis vector $\langle \overrightarrow{x}|$, we obtain equivalent Schrodinger equation in coordinate representation:

$$i\hbar \frac{d}{dt} \Psi(\overrightarrow{x}, t) = \hat{H}(\overrightarrow{x}) \Psi(\overrightarrow{x}, t).$$

While an expert will necessarily regard Eqs.(1-3) as a great simplification when thinking of the content of quantum physics, the novice often understandably reels under the weight of the immense abstraction. We learn much about student thinking from from the answers given by our best students. For example, we find a common error when studying 1D quantum mechanics is a student treating $\Psi(x)$ and $|\Psi\rangle$ interchangeably, ignoring the fact that the first is a scalar but the ket corresponds to a column vector. For example, they may write incorrectly

(4)
$$\langle p|\Psi\rangle|x\rangle = |\Psi\rangle\langle p|x\rangle$$
 (incorrect!)

or some similar abberation. To avoid these types of misconceptions, a number of educators and textbook authors have stressed incorporating a numerical calculation aspect to quantum courses [4,5,6,7,8,9]. The motive is simple. Anyone who has done numerical calculations can't help but regard a ket $|\Psi\rangle$ as a column vector, a bra $\langle\Psi|$ as a row vector and an operator H as a matrix because that is how they concretely represented in the computer. Introducing a computational aspect to the course provides one further benefit: it gives the beginning quantum student the sense that he or she is being empowered to solve real problems that may not have simple, analytic solutions.

With these motivations in mind, we have developed MATLAB codes [10] for solving typical 1D problems found in the first part of a junior level quantum course based on Griffith's book [11]. We chose MATLAB for our

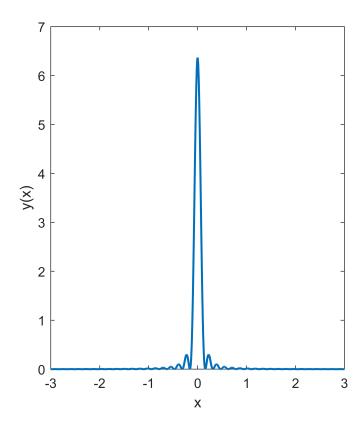
programming environment because the MATLAB syntax is especially simple for the typical matrix operations used in 1D quantum mechanics problems and because of the ease of plotting functions. While some MATLAB numerical recipes have previously been published by others, [12,13] the exercises we share here are special because they emphasize simplicity and quantum pedagogy, not numerical efficiency. Our point has been to provide exercises which show students how to numerically solve 1D problems in such a way that emphasizes the column vector aspect of kets, the row vector aspect of bras and the matrix aspect of operators. Exercises using more efficient MATLAB ODE solvers or finite-element techniques are omitted because they do not serve this immediate purpose. In part II of this article, we hope to share MATLAB codes which can be used in conjunction with teaching topics pertaining to angular momentum and non-commuting observables.

1. Functions as Vectors

To start students thinking of functions as column vector-like objects, it is very useful to introduce them to plotting and integrating functions in the MATLAB environment. Interestingly enough, the plot command in MATLAB takes vectors as its basic input element. As shown in Program 1 below, to plot a function f(x) in MATLAB, we first generate two vectors: a vector of x values and a vector of y values where y = f(x). The command plot(x,y,'r') then generates a plot window containing the points (x_i,y_i) displayed as red points ('r'). Having specified both vectors x and y, to evaluate the integral $\int_{-L}^{L} y \, dx$, we need only sum all the y values and multiply by dx.

```
%Program 1: Numerical Integration and Plotting using MATLAB
N = 1e6;
                         % No. of points.
L = 500;
                         % Range of x: from -L to L.
x = linspace(-L, L, N).'; % Generate column vector with N
                         % x values ranging from -L to L.
% Alternative Trial functions:
% To select one, delete the comment character '%' at the beginning.
%y = exp(-x.^2/16);
                                     % Gaussian centered at x=0.
y = ((2/pi)^0.5) * exp(-2*(x-1).^2); % Normed Gaussian at x=1.
%y = (1-x.^2).^{-1};
                                     % Symmetric fcn, blows up at x=1.
%y = (cos(pi*x)).^2;
                                     % Cosine fcn.
%y = exp(i*pi*x);
                                     % Complex exponential.
y = \sin(pi*x/L) \cdot \cos(pi*x/L); % Product of sinx times cosx.
% Element-wise product .* of two row vectors is a row vector as well
%y = sin(pi*x/L) .* sin(pi*x/L);
                                     % Product of sin(nx) times sin(mx).
%A = 100; y = \sin(x*A) ./ (pi*x);
                                     % Rep. of delta fcn.
A = 20; y = (\sin(x*A).^2) ./ (pi*(A*x.^2)); % Another rep. of delta fcn.
% Observe: a function y(x) is numerically represented by a vector!
% Plot a vector/function.
axis([-3 \ 3 \ 0 \ 7]); \% Optimized for <math>sin(x).^2 ./ (pi * x.^2).
% plot(x,real(y), 'r', x, imag(y), 'b'); % Plots real&imag y vs x.
                                        % Optimized for sin(x) ./ (pi * x).
% axis([-2 2 -8 40]);
xlabel('x');
```

```
ylabel('y(x)');
daspect([1 1 1]);
```



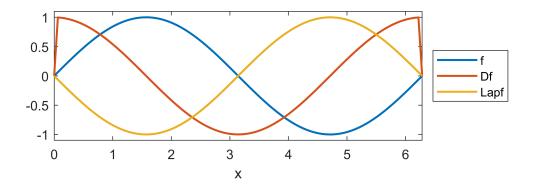
2. Differential Operators as Matrices

Just as wavefunction in real-space representation is represented by a column vector $|f\rangle = \sum_j f(x_j)|x_j\rangle$ in the computer, for numerical purposes an operator \widehat{D} acting on $|f\rangle$ (gradient, Laplacian, etc.) is represented by a matrix D that acts on column corresponding to realspace representation of $|f\rangle$. As illustrated in Program 2, MATLAB provides many useful, intuitive, well-documented commands for generating easily and efficiently matrices D that correspond to basic operators encountered in quantum mechanics \widehat{D} [10]. Two examples are the commands ones and diag. The command ones (a,b) generates an $a \times b$ matrix of ones. The command diag(A,n) generates a matrix with the elements of the vector A placed along the nth diagonal and zeros everywhere else. The central diagonal corresponds to n=0, the diagonal above/ below the central one corresponds to n=1/-1, etc....

An exercise we suggest is for students to verify that the derivative matrix is not Hermitian while the derivative matrix times the imaginary number i is. This can be very valuable for promoting student understanding if done in conjunction with the proof usually given for the differential operator.

```
%-----
% Program 2: Calculate first and second derivative numerically
% showing how to write differential operator as a matrix.
%------
% Parameters for solving problem in the interval 0 < x < L.</pre>
```

```
L = 2*pi;
                        % Interval Length.
                        % No. of coordinate points.
N = 100;
x = linspace(0, L, N).'; % Coordinate vector.
dx = x(2) - x(1);
                   % Coordinate step.
% Two-point finite-difference representation of Derivative.
D = (diag(ones((N-1),1),1) - diag(ones((N-1),1),-1)) / (2*dx);
% Next modify D so that it is consistent with f(0) = f(L) = 0.
D(1, 1) = 0; D(1, 2) = 0; D(2, 1) = 0;
D(N, N-1) = 0; D(N-1, N) = 0; D(N, N) = 0;
%Laplacian can be obtained directly by D*D, but better recipe is to
%consider a three-point finite-difference representation of Laplacian
Lap = (diag(ones((N-1),1),-1) - 2*diag(ones(N,1),0) ...
     + diag(ones((N-1),1), 1)) / (dx^2);
% Next modify Lap so that it is consistent with f(0) = f(L) = 0.
Lap(1, 1) = 0; Lap(1, 2) = 0; Lap(2, 1) = 0;
Lap(N, N-1) = 0; Lap(N-1, N) = 0; Lap(N, N) = 0;
% To verify that D*f corresponds to taking the derivative of f
% and Lap*f corresponds to taking a second derviative of f,
% define f = \sin(x) or choose your own f that vanishes at boundaries,
f = sin(x); %f is a column vector, as x is a column
% and try the following:
Df = D * f;
Lapf = Lap * f;
plot(x,
         f, ...
     x,Lapf); %instead of three-point Laplacian D*D works as well
legend('f','Df','Lapf','Location','EastOutside');
xlabel('x');
axis([0 2*pi -1.1 1.1]);
daspect([1 1 1]);
```



3. Infinite Square Well

When solving Eq. (1), the method of separation of variables entails that as an intermediate step we look for the separable solutions (separable means of the form: $|\psi\rangle = T(t) * X(x)$).

First we note that Shrodinger equation (1) can be formally solved as: $|\Psi(t)\rangle = \exp(-i\hat{H}t/\hbar)|\Psi(0)\rangle$.

Let's assume that at t=0 state of the system coinsides with one of the energy eigenstates $|\Psi(0)\rangle = |\Psi_E\rangle$. Then, at time t, we will have:

(5)
$$|\Psi(t)\rangle = |\Psi_E\rangle \exp(-iEt/\hbar)$$
,

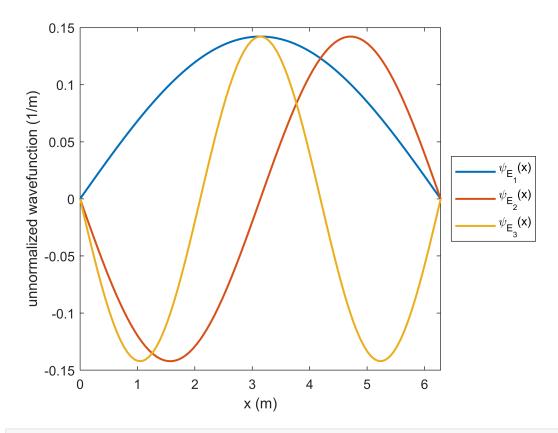
where $|\Psi_{\it E}\rangle$, as an energy operator basis state satisfies the time-independent Schrodinger equation

(6)
$$\hat{H} |\Psi_E\rangle = E |\Psi_E\rangle$$
.

In solving Eq. (6) we are solving for the eigenvalues E and eigenvectors $|\Psi_E\rangle$ of \hat{H} . In MATLAB, the command [V,E]=eig(H) does precisely this: it generates two matrices. The first matrix V has as its columns the eigenvectors $|\Psi_E\rangle$ (in general not normalized to unity). The second matrix E is a diagonal matrix with the eigenvalues E_i corresponding to the eigenvectors $|\Psi_{E_i}\rangle$ placed along the central diagonal. If columns of V are normalized eigenstates, then E=V'*H*V, where prime means hermit conjugation (transposition + complex conjugation).

We can use the command E = diag(E) to convert diagonal matrix into a column vector. In Program 3, we solve for the eigenfunctions and eigenvalues for the infinite square well Hamiltonian. For brevity, we omit the commands setting the parameters L, N, x, and dx.

```
% Program 3: Matrix representation of differential operators,
% Solving for Eigenvectors & Eigenvalues of Infinite Square Well.
% For brevity we omit the commands setting the parameters L, N,
% x and dx. We also omit the commands defining the matrix Lap.
% These would be the same as in Program 2 above.
% Total Hamiltonian where hbar=1 and m=1.
hbar = 1;
m = 1;
H = -(1/2)*(hbar^2/m)*Lap;
% Solve for eigenvector matrix V and eigenvalue matrix E of H.
[V, E] = eig(H);
% Plot lowest 3 eigenfunctions.
plot(x,V(:,3), x,V(:,4), x,V(:,5)); %V(:,1) and V(:,2) are boundary zero-energy
%nonphysical states, they do not vanish at boundaries and exist due
%to our convention for Laplacian at edges.
legend('\psi_{E_1}(x)','\psi_{E_2}(x)','\psi_{E_3}(x)','Location','EastOutside');
xlabel('x (m)');
ylabel('unnormalized wavefunction (1/m)');
ax = gca; % Get the Current Axes object
ax.XLim = [0 2*pi];
```



diag(E) % display a column vector containing the eigenvalues.

```
ans = 100×1
0
0
0.1250
0.4998
1.1242
1.9973
3.1184
4.4864
6.0999
7.9571
```

Note that in the MATLAB syntax the object V(:,3) specifies the column vector consisting of all the elements in column 3 of matrix V. Similarly V(2,:) is the row vector consisting of all elements in row 2 of V; V(3,1) is the element at row 3, column 1 of V; and V(2,1:3) is a 3-element row vector consisting of elements V(2,1), V(2,2) and V(2,3).

4. Arbitrary Potentials

Numerical solution of Eq. (1) is not limited to any particular potential. Program 4 below gives example MATLAB codes solving the time independent Schrodinger equation for finite square well potentials, the harmonic oscillator potential and even for potentials that can only solved numerically such as the quartic potential $U = x^4$. In order to minimize the amount of RAM required, the codes shown make use of sparse matrices, where only the non-zero elements of the matrices are stored. The commands for sparse matrices are very similar to those

for non-sparse matrices. For example, the command [V,E] = eigs(H,nmodes) provides the nmodes lowest energy eigenvectors V of of the sparse matrix H.

```
% Program 4: Find several lowest eigenmodes V(x) and
% eigenenergies E of 1D Schrodinger equation
      -1/2*hbar^2/m(d2/dx2)V(x) + U(x)V(x) = EV(x)
% for arbitrary potentials U(x).
% Parameters for solving problem in the interval -L < x < L.
% PARAMETERS:
L = 5;
                        % Interval Length.
N = 1000;
                       % No of points.
x = linspace(-L, L, N).';% Coordinate vector.
dx = x(2) - x(1); % Coordinate step.
% POTENTIAL, choose one or make your own.
U = 1/2*100*x.^(2); % quadratic harmonic oscillator potential
%U = 1/2*x.^{(4)}; % quartic potential
% Finite square well of width 2w and depth given.
%w = L/50;
%U = -500*(heaviside(x+w)-heaviside(x-w));
% Two finite square wells of width 2w and distance 2a apart.
%w = L/50;
%a = 3*w;
%U = -200*(heaviside(x+w-a) - heaviside(x-w-a) ...
% + heaviside(x+w+a) - heaviside(x-w+a));
% Three-point finite-difference representation of Laplacian
% using sparse matrices, where you save memory by only
% storing non-zero matrix elements.
e = ones(N,1); % a column of ones
Lap = spdiags([e -2*e e], [-1 0 1], N, N) / dx^2;
% put -2e on the main diagonal and e-s on upper and lower diagonals
% Total Hamiltonian.
hbar = 1;
m = 1;
H = -1/2*(hbar^2/m)*Lap + spdiags(U,0,N,N); % 0 indicates main diagonal
% put vector U on the main diagonal of NxN sparse matrix
% Find lowest nmodes eigenvectors and eigenvalues of sparse matrix.
nmodes = 3;
[V,E] = eigs(H,nmodes,'smallestreal'); % find eigs.
[E,ind] = sort(diag(E));
                                  % convert E to vector and sort low to high.
V = V(:,ind);
                                   % rearrange corresponding eigenvectors.
% Generate plot of lowest energy eigenvectors V(x) and U(x).
Usc = 10 * U * max(abs(V(:))) / max(abs(U));% rescale U for plotting.
                         \% plot V(x) and rescaled U(x).
plot(x,V, x,Usc,'--k');
xlabel('x (m)');
ylabel('unnormalized wavefunction');
```

```
% Add legend showing Energy of plotted V(x).
legendLabels = [repmat('E = ',nmodes,1), num2str(E)];
legend(legendLabels) % place lengend string on plot.
ax = gca;
ax.XLim = sqrt(2)*[-1 1];
```

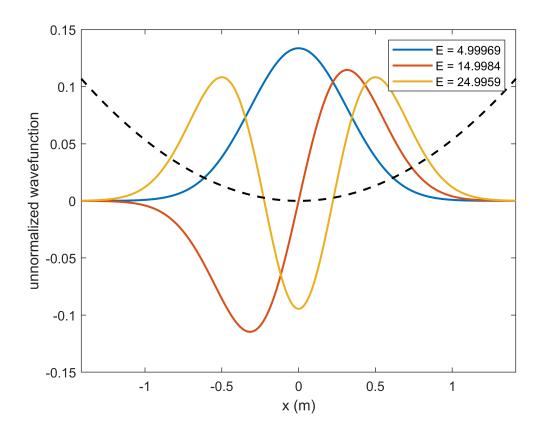


Fig. 1 shows the plot obtained from Program 4 for the potential $U = \frac{1}{2} \cdot 100 \cdot x^2$. Note that the 3 lowest energies displayed in the figure are just as expected due to the analytic formula

$$(7) \quad E = \hbar\omega \left(n + \frac{1}{2}\right)$$

with n= integer and $\omega=\sqrt{\frac{k}{m}}=10\,\mathrm{rad/s}$.

5. Time Dependent Phenomena

The separable solutions $|\Psi_E(t)\rangle$ are only a subset of all possible solutions of Eq. (1). Fortunately, they are complete set (time independent hamiltonian is hermitian operator and has a complete set of eigenstates, which are separabe) so that we can construct the general solution *via* the linear superposition

$$(8) \quad |\Psi(t)\rangle = \sum_{E} |\Psi_{E}\rangle\langle\Psi_{E}|\Psi(t)\rangle = \sum_{E} a_{E}|\Psi_{E}\rangle \exp(-iEt/\hbar),$$

where $a_E = \langle \Psi_E | \Psi(0) \rangle$ are projection of the generic (non-separable) state at t=0 onto the energy Eigenstates and the sum is over all possible values of E. The important difference between the separable solutions (5) and the general solution (8) is that the probability densities derived from the general solutions are time-dependent whereas those derived from the separable solutions are not. A very apt demonstration of this is provided in the Program 5 which calculates the time-dependent probability density $\rho(x,t)$ for a particle trapped in a pair of finite-square wells whose initial state $|\Psi(0)\rangle$ is set equal to the the equally-weighted superposition

(9)
$$|\Psi(0)\rangle = \frac{1}{\sqrt{2}} (|\Psi_{E_0}\rangle + |\Psi_{E_1}\rangle)$$

of the ground state $|\Psi_{E_0}\rangle$ and the first excited state $|\Psi_{E_1}\rangle$ of the double well system. As snapshots of the program output show in Fig. 2, the particle is initially completely localized in the rightmost well. However, due to $E_0 \neq E_1$, the probability density

(10)
$$\rho(x,t) = \frac{1}{2} \left[|\psi_{E_0}(x)|^2 + |\psi_{E_1}(x)|^2 + 2\psi_{E_0}(x)\psi_{E_1}(x)\cos((E_1 - E_0)t/\hbar) \right]$$

is time-dependent, oscillating between the $\rho(x,t)=\rho_R(x)$ for $t=2\pi n\hbar/(E_1-E_0)$ and n=0,1,2,3,... that corresponds to the particle being entirely in the right well

(11)
$$\rho_R(x) = |\psi_{E_0}(x) + \psi_{E_1}(x)|^2$$

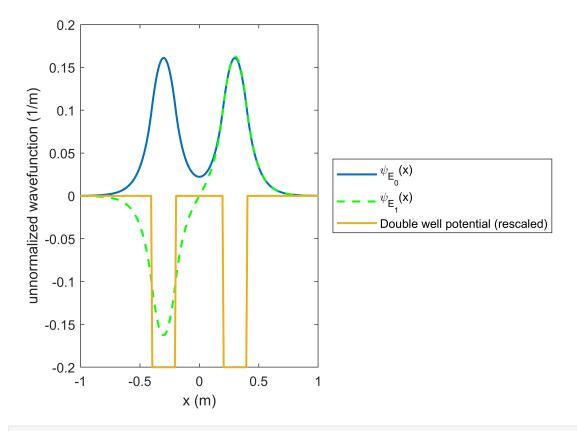
and $\rho(x,t)=\rho_L(x)$ for $t=2\pi(n+1/2)\hbar/(E_1-E_0)$ and n=0,1,2,3,... corresponding to the particle being entirely in the left well

(12)
$$\rho_L(x) = |\psi_{E_0}(x) - \psi_{E_1}(x)|^2$$
.

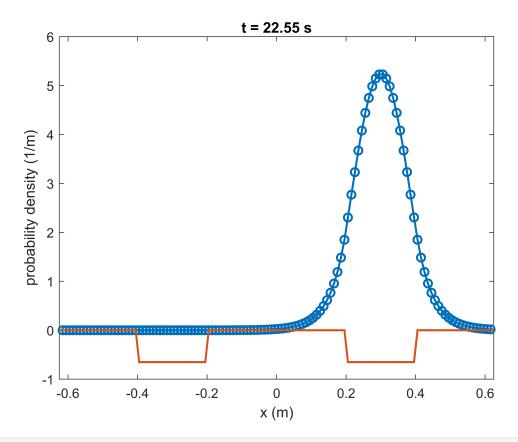
In deriving Eq. (10) we used the fact that both ψ_{E_0} and ψ_{E_1} are real functions.

By observing the period with which $\rho(x,t)$ oscillates in the simulation output shown in Fig. 2 students can verify that it is the same as the period of oscillation $2\pi\hbar/(E_1-E_0)$ expected from Eq. (10).

```
H = -(1/2)*Lap + spdiags(U,0,N,N);
% Find and sort lowest nmodes eigenvectors and eigenvalues of H.
nmodes = 2;
[V,E] = eigs(H,nmodes,'smallestreal');
[E,ind] = sort(diag(E)); % Convert E to vector and sort low to high.
                          % Rearrange coresponding eigenvectors, usually they are not normalized
        = V(:,ind);
% Rescale eigenvectors so that they are always
% positive at the center of the right well.
% Still they may not be normalized to unity.
for c = 1:nmodes
    V(:,c) = V(:,c) / sign(V((3*N/4),c));
% Compute and display normalized prob. density rho(x,t).
\% Parameters for solving the problem in the interval 0 < t < TF.
TF = 4*pi*hbar/(E(2)-E(1)); % Length of time interval.
NT = 100;
                            % No. of time points.
t = linspace(0,TF,NT);
                           % Time vector.
% Compute probability normalization constant (at T=0).
psi_o = 0.5*V(:,1) + 0.5*V(:,2); % Wavefunction at T=0. 
 sq_norm = psi_o' * psi_o * dx; % Square norm = |\langle ff|ff \rangle|^2.
Usc = 4*U*max(abs(V(:))) / max(abs(U)); % Rescale U for plotting.
% Compute and display rho(x,t) for each time t.
% Plot lowest 2 eigenfunctions.
plot(x,V(:,1), x,V(:,2),'--g',x,U/500);
legend('\psi_{E_0}(x)','\psi_{E_1}(x)','Double well potential (rescaled)','Location','EastOuts
xlabel('x (m)');
ylabel('unnormalized wavefunction (1/m)');
ax = gca; % Get the Current Axes object
ax.XLim = [-1 1];
```



```
vw1 = VideoWriter('doubleWell.avi'); % Prepare the new file.
% vw1.FrameRate=30 by default
open(vw1);
for jj = 1:NT % time index parameter for stepping through loop.
    % Compute wavefunction psi(x,t) and rho(x,t) at t=t(jj).
    psi = 0.5*V(:,1)*exp(-1i*E(1)*t(jj)) ...
       + 0.5*V(:,2)*exp(-1i*E(2)*t(jj));
    rho = conj(psi) .* psi / sq_norm; % Normalized probability density.
   % Plot rho(x,jj) and rescaled potential energy Usc.
    plot(x,rho,'o-', x, Usc,'.-');
    axis([-L/8 L/8 -1 6]);
    xlabel('x (m)');
    ylabel('probability density (1/m)');
    title(['t = ' num2str(t(jj), '%05.2f') ' s']);
    writeVideo(vw1, getframe(gcf));
    % getframe returns a movie frame, a snapshot (pixmap) of the current axes or figure.
end
```



close(vw1);
%implay('doubleWell.avi')

6. Wavepackets and Step Potentials

Wavepackets are another time-dependent phenomenon encountered in undergraduate quantum mechanics for which numerical solution techniques have been typically advocated in the hopes of promoting intuitive acceptance and understanding of approximations necessarily invoked in more formal, analytic treatments. Program 6 calculates and displays the time evolution of a wavepacket for one of two possible potentials, either U=0 or a step potential $U=U_0\Theta(x-L)$. The initial wavepacket is generated as the Fast Fourier Transform of a Gaussian momentum distribution centered on a particular value of the wavevector k_0 . Because the wavepacket is composed of a distribution of different k_0 , different parts of the wavepacket move with different speeds, which leads to the wave packet spreading out in space as it moves.

While there is a distribution of velocities within the wavepacket, two velocities in particular are useful in characterizing it. The phase velocity $v_p = \omega/k = E/p = \hbar k_0/2m$ is the velocity of the plane wave component which has wavevector k_0 . The group velocity $v_g = \hbar k_0/m$ is the velocity with which the expectation value $\langle x \rangle$ moves and is the same as the classical particle velocity associated with the momentum $p = \hbar k$. Choosing U = 0, students can modify this program to plot $\langle x \rangle$ vs t. They can extract the group velocity from their numerical simulation and observe that indeed $v_g = 2v_p$ for a typical wave packet. Students can also observe

that, while v_g matches the particle speed from classical mechanics, the wavepacket spreads out as time elapses.

In Program 6, we propagate the wave function forward via the formal solution

(13)
$$|\Psi(t)\rangle = \exp(-i\hat{H}t/\hbar)|\Psi(0)\rangle$$
,

where the matrix H corresponding to Hamiltonian \widehat{H} is in the exponential. This solution is equivalent to Eq. (6), as can be shown by simple substitution. Moreover, MATLAB has no trouble exponentiating the matrix that numerically representing the Hamiltonian operator as long as the matrix is small enough to fit in the available computer memory.

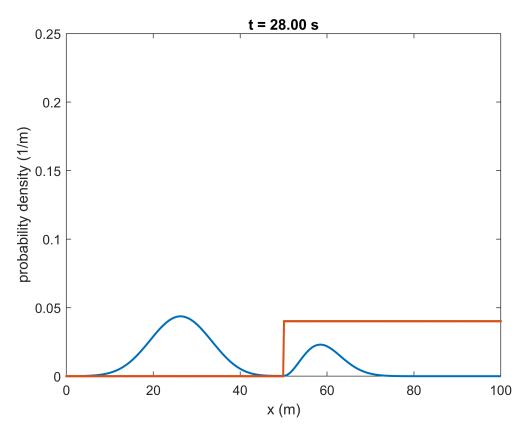
Even more interestingly, students can use this method to investigate scattering of wavepackets from various potentials, including the step potential $U=U_0\Theta(x-L/2)$. In Fig. 3, we show the results of what happens as the wavepacket impinges on the potential barrier. The parameters characterizing the initial wavepacket have been deliberately chosen so that the wings do not fall outside the simulation area and initially also do not overlap the barrier on the right. If $\langle E \rangle \ll U_0$, the wavepacket is completely reflected from the barrier. If $\langle E \rangle \approx U_0$, a portion of the wave is reflected and a portion is transmitted through. If $\langle E \rangle \gg U_0$, almost all of the wave is transmitted.

In Fig. 4 we compare the reflection probability R calculated numerically using Program 6 with R calculated by averaging the single-mode[11] expression

(14)
$$R(E) = \left| \frac{\sqrt{E} - \sqrt{E - U_0}}{\sqrt{E} + \sqrt{E - U_0}} \right|^2$$

over the distribution of energies in the initial wavepacket. While the numerically and analytically estimated R are found to agree for large and small $\langle E \rangle/U_0$, there is a noticeable discrepancy due to the shortcomings of the numerical simulation for $\langle E \rangle/U_0 \approx 1$. This discrepancy can be reduced significantly by increasing the number of points in the simulation to 1600 but only at the cost of significantly slowing down the speed of the computation. For our purposes, the importance comparing the analytical and numerical calculations is that it gives students a baseline from which to form an opinion or intuition regarding the accuracy of Eq. (14).

```
% Make psi(x,0) from Gaussian kspace wavefunction phi(k) using
% fast fourier transform.
phi = exp(-a*(k-ko).^2).* exp(-1i*6*k.^2); % Unnormalized phi(k).
psi = ifft(phi);
                                  % Multiplies phi by exp(ikx) and integrates vs x.
psi = psi / sqrt(psi' * psi * dx); % Normalize psi(x,0).
% Expectation value of energy; e.g. for the parameters
% chosen above \langle E \rangle = 2.062.
avgE = phi' * 0.5*diag(k.^2,0) * phi * dk / (phi' * phi * dk);
% CHOOSE POTENTIAL U(X): Either U = \emptyset OR
% U = step potential of height avgE that is located at x = L/2.
%U = 0*heaviside(x-(L/2)); % Free particle wave packet evolution.
U = avgE*heaviside(x-(L/2)); % Scattering off step potential.
% Finite-difference representation of Laplacian and Hamiltonian.
e = ones(N,1);
Lap = spdiags([e -2*e e],[-1 0 1],N,N) / dx^2;
H = -(1/2)*Lap + spdiags(U,0,N,N);
\% Parameters for computing psi(x,t) at different times 0 < t < TF.
NT = 200;
                      % No. of time steps.
TF = 28;
t = linspace(0,TF,NT); % time vector
dt = t(2) - t(1); % time step
hbar = 1;
% time displacement operator.
Evolution_Operator = expm(-1i * H * dt / hbar);
% Simulate rho(x,t) and plot for each t
%-----
vw2 = VideoWriter('propagation.avi');
vw2.FrameRate=35;
open(vw2);
for jj = 1:NT % Time index for loop.
    % Calculate probability density rho(x,t).
    psi = Evolution Operator * psi;  % Calculate psi(t+\Delta t) from psi(t)
    rho = conj(psi) .* psi; % rho(x,t).
    plot(x,rho,x,U/50,'.-'); %potential is rescaled to fit nicely in the plot
    axis([0 L 0 0.25]);
    xlabel('x (m)');
   ylabel('probability density (1/m)');
   title(['t = ' num2str(t(jj), '%05.2f') ' s']);
   writeVideo(vw2, getframe(gcf));
end
```



```
close(vw2);
%implay('propagation.avi') % uncomment this to launch automatically immage
%player
% Calculate Reflection probability
R = 0;
for a = 1:floor(N/2)
    R = R + rho(a); % density at the eft of step barrier
end
R = R * dx
```

A Note on Units in Our Programs

R = 0.7501

When doing numerical calculations, it is important to minimize the effect of rounding errors by choosing units such that the variables used in the simulation are of the order of unity. In the programs presented here, our focus being undergraduate physics students, we wanted to avoid unnecessarily complicating matters. To make the equations more familiar to the students, we explicitly left constants such as \hbar in the formulas and chose units such that $\hbar=1$ and m=1. We recognize that others may have other opinions on how to address this issue. An alternative approach used in research is to recast the equations in terms of dimensionless variables, for example by rescaling the energy to make it dimensionless by expressing it in terms of some characteristic energy in the problem. In a more advanced course for graduate students or in a course in numerical methods, such is an approach which would be preferable.

Conclusions

One benefit of incorporating numerical simulation into the teaching of quantum mechanics, as we have mentioned, is the development of student intuition. Another is showing students that non-ideal, real-world problems can be solved using the concepts they learn in the classroom. However, our experimentation incorporating these simulations in quantum physics at WPI during the past year has shown us that the most important benefit is a type of side-effect to doing numerical simulation: the acceptance on an intuitive level by the student that functions are like vectors and differential operators are like matrices. While in the present paper, we have only had sufficient space to present the most illustrative MATLAB codes, our goal is to eventually make available a more complete set of polished codes available for downloading either from the authors or directly from the file exchange at MATLAB Central [14].

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