# **Lesson 4: Remote Sensing Mars**

### **Summary**

This learning module and related laboratory exercise exposes students to remote sensing techniques utilized on Mars.

### **Learning Goals**

#### Students will be able to:

- Apply the concepts of scale and context in remote sensing imagery.
- View THEMIS and HiRISE images and interpret major geomorphic features using Google Mars and associated homework activities.
- Understand how MOLA generates its image data by applying the fundamental equations in an experiment.

#### **Context for Use**

This learning module is meant for adaptation in an introductory earth science course and/or planetary science course. It is advised that the teacher compare Earth-based remote sensing instrumentation for context/reference such as LandSat 7.

### **Description and Teaching Materials**

In-Class Activity

In-Class Activity 1: Scale and Context In-Class Activity 2: MOLA simulation *Homework/Lab* 

Homework 1: Google Mars-Following Opportunity

Homework 2: Mars Image Analysis

#### **Teaching Notes and Tips**

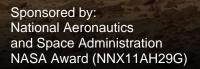
- 1. The *In-Class Activities* can be utilized as homework as well. Students will have a lab-write up associated with the *MOLA simulation*.
- 2. For a large class size >20 you may either have a separate lab time/class for different sections or demonstrate the lab with the entire class and employ student participation.

- 3. We advise instructors to compare Earthbased remote sensing packages such as Landsat 7 for context.
- 4. In preparation for the MOLA simulation *In-Class Activity* instructors must gather a few materials (see the *MOLA simulation* for further clarification).

#### Assessment

- The MOLA simulation Lab write-up will assess the student's understanding of the MOLA instrument and MOLA's utility.
- The Google Mars homework will assess whether or not students can successfully navigate the Google Mars software and begin to interpret the data provided by Google Mars.







#### **References and Resources**

- 1. THEMIS images url: <a href="http://themis.asu.edu/">http://themis.asu.edu/</a>
- 2. LANDSAT 7 images url: <a href="http://landsat.gsfc.nasa.gov/images/">http://landsat.gsfc.nasa.gov/images/</a>
- 3. HiRISE 13 April 2011 YouTube video: <a href="http://www.youtube.com/watch?v=-U6-uYDtuSg">http://www.youtube.com/watch?v=-U6-uYDtuSg</a>
- 4. MRO/HiRISE All HiClips revisited (Feb 2012) YouTube Video: http://www.youtube.com/watch?v=YVDUQjJbjyc
- 5. MOLA images url: http://mola.gsfc.nasa.gov/index.html
- 6. Ping-Pong Lab (NASA): <a href="http://mola.gsfc.nasa.gov/pingpong.html">http://mola.gsfc.nasa.gov/pingpong.html</a>



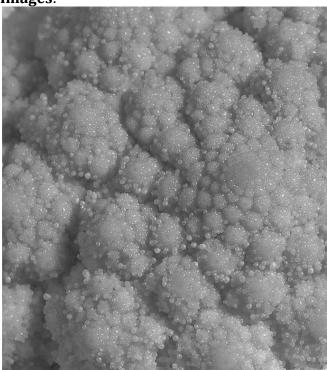
In-Class Activity 1
Remote Sensing Mars MFE

Scale and Context

**Purpose**: Recognize the purpose and need for understanding the scale and context of various remote sensing imaging techniques.

**Preparation**: Print images (or share in .ppt, see *Image File*) without their captions.

Images:



**Figure 1:** Close-up of broccoli romanesco Image credit: Petr Kratochvil, public domain.

### **Context Image**







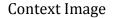
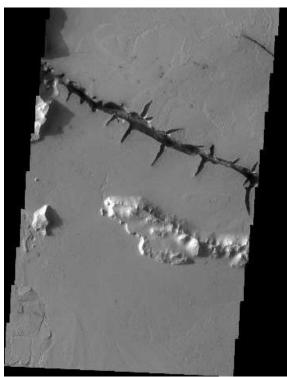


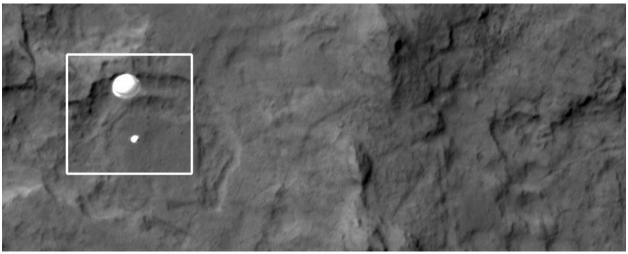


Figure 2: Close-up of a pine cone. Image credit: Petr Kratochvil, public domain



**Figure 3** THEMIS Image #V13300013, Image width: 25km; Lat 7.3/Long 161.3. Image credit: NASA/JPL/ASU; Image Source: <a href="http://themis.asu.edu/zoom-20050225a">http://themis.asu.edu/zoom-20050225a</a>





**Figure 4:** HiRISE image (ESP\_028256\_9022) of Curiosity descending to the Martian surface acquired August 5, 2012. Scale = 33.6 cm/pixel; Image Source: <a href="http://hirise.lpl.arizona.edu/releases/msl-descent.php">http://hirise.lpl.arizona.edu/releases/msl-descent.php</a>

### **Engage**

Referring to Figures 1-2 ask students to:

- 1. Hypothesize what is pictured in Figure 1 & 2.
- 2. Describe the basis for their hypothesis.
- 3. List at least (4) ideas for gathering other information that could aid in determining what the image is exactly. Explain their reasoning for their choices.

### **Explore**

Time to apply these terms for images of Mars

- 1. Have students interpret Figures 1 & 2. Then, define terms of scale and context, and apply knowledge to Mars imagery.
- 2. Encourage students to interpret Figure 3 & 4. Provide students with the scale and context for the images (given in this version in the image caption).
- 3. Ask students to discuss whether or not knowing the *scale* and *context* of the images has aided them in interpretation.

### **Explain**

- 1. Discuss and determine definitions of the terms *scale* and *context*.
- 2. Reveal the identity of the objects shown in Figure 1 & 2 (knowing what these are given in the image captions, the scales of these are small cm to mm).
- 3. Ask students to provide a *scale* and *context* for each image.



#### **Elaborate**

- 1. Time permitting, have students explore mission websites such as THEMIS (<a href="http://themis.asu.edu">http://themis.asu.edu</a>) or HiRISE (<a href="http://hirise.lpl.arizona.edu">http://hirise.lpl.arizona.edu</a>)
- 2. Have them choose an image that interests them. What caught their interest? What features do they see?
- 3. Can they find the scale and context of the image? After knowing the scale and context, does their interpretation of the image change?

#### **Evaluate**

1. Evaluate student responses as they interpret the images. Each student should select an image and submit an explanation of the scale and context for evaluation. Do students have a clear understanding of scale and context as indicated by their answers in the *Elaboration* section?



In-Class Activity 2
Remote Sensing\_MFE

MOLA Simulation\*

**Purpose**: Understand how we explore the surface of Mars via remote sensing techniques by performing a ping-pong experiment.

**Preparation**: This experiment requires some space (like a hallway, or space near a cleared wall in a classroom) and will take some prep time ~ 15 mins., and ~ 20-30 mins. for students to perform the exercise. Some of the data plotting can be done as homework.

#### Materials Needed:

Masking tape, meter sticks or measuring tape, ping pong balls, stopwatch or watch timer, bricks or blocks that allow ping pong balls to bounce (textbooks are ineffective)

### **Engage**

How do we know what the surface of Mars is like, especially for areas that we have only seen from a distance? Think about how dolphins know the difference between a BB gun pellet and a kernel of corn from 50' (or 16 m away). Could a similar type of detection be used to decipher the surface of Mars?

Ref. <a href="http://science.howstuffworks.com/zoology/marine-life/dolphin-disarm-seamine1.htm">http://science.howstuffworks.com/zoology/marine-life/dolphin-disarm-seamine1.htm</a>

#### **Explore**

Perform a ping-pong experiment.

#### Procedure:

The students must have at least 2 people in their group (3 per group is preferable).

#### Step 1:

- 1. Place 2 strips of tape on the wall, one approximately 2 meters (200 cm) high and the other 45 cm high. Both should be at least 200 cm long and parallel; students will be using these as the points to start and stop the stopwatch.
- 2. One partner should hold the ping-pong ball between the first finger and thumb next to the higher piece of tape approximately one inch from the wall.
- 3. One partner, the "timer", should have a stopwatch and have his/her eyes level with the second piece of tape. A third partner, if available, should record the results of each ball drop using the attached data sheet. \*Note: Use a spreadsheet for recording and calculating the data.
- 4. Drop the ball. Start the stop watch as soon as the ball begins to drop.



- 5. The timer will stop the watch when the ball rebounds and reaches the lower line, i.e. the clock starts when the ball drops and stops when the ball reaches the second piece of tape. Record the time on the data sheet. Repeat this step four more times.
- 6. Calculate the velocities (V=D/T). The distance (D) is the combination of the height of the high tape plus the height of the low tape. After finding the velocity for each of the trials, find the average velocity of the ping-pong ball. This average will be used later in this lab as the students' baseline for comparing data. Ask the students if for each trial are they measuring an average or instantaneous velocity?
- 7. Many spacecraft use lasers (light) to determine topography similar to how the students are using a ping pong ball. However there is a potential over-simplification in using a ping pong ball as an analog to a laser. What are the issues? (Hint: Think about velocity vs. acceleration.)

Data Table I (Baseline, datum)

Drop	Distance Ball Traveled	Time (Seconds)	Velocity (distance/time)
1			
2			
3			
4			
5			
		Average Velocity	

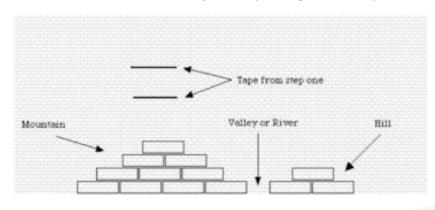
### Step 2:

Now that the students have found the velocity of the ping-pong ball, they will use this information to plot the topography of a transect along the surface of an imaginary asteroid. They will be creating your own asteroid terrain on the floor against the wall where they just did Step 1.

1. Create the topography model of an asteroid along the wall where the students did Step One. In order to do this, they need to place wooden blocks *against* the wall in a line about 6 feet long. Be sure that they build some hills, mountains, valleys, etc. (see Figure below).

Ping-pong experiment layout diagram.

2. Starting at the beginning of the top piece of tape, place a mark every 20cm. The bottom piece does not need to be marked. Measure and record the topography heights in the far right column of Table II as a check.





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- 3. Again, starting at the 0 cm mark the students made at the beginning of the tape, they will drop the ping-pong ball as they did in Step One, and record the time in Data Table II. Drop the ball every 20cm along the tape until they reach the end. Be sure to be as accurate as possible with the timing.
- 4. Find the average time for each of the intervals and record it in the data table.
- 5. Have the students exchange their average time data with another student group (Table II) of just the cm interval, the average time. They will interpret each other's data to see if they can identify the topography of the other group's asteroid.
- 6. Next, fill in the distance traveled in Table II by multiplying the average velocity from Table I by the average time just calculated at each interval.
- 7. In Data Table III, take the original distance traveled (height of high tape plus height of low tape, which will be the same for every interval) and fill in the first column of the table.
- 8. In the second column, take the distance the ball traveled from the column in Table II (last column on right) and copy that information to the  $2^{nd}$  column of Table III. Now, for the last column, simply subtract the 1st column data from the  $2^{nd}$  column data (the difference between the two) to determine the altitude of their modeled topography.
- 9. Plot the data (with interval/distance on the x-axis and altitude on the y-axis). Connect the dots to create their transect. Does their image match the true topography? If not, explain why it is different.



#### **Data Table II**

					Tuble II	1
Interval	Trial 1	Trial 2	Trial 3	Time	Distance Ball Traveled =	Known measured height
				Average	(velocity*average time)	of placed block
					(cm)	topography (cm)
0 cm						
20 cm						
40 cm						
60 cm						
80 cm						
100 cm						
120 cm						
140 cm						
160 cm						
180 cm						
200 cm						

----- tear here to give Table III to blind student group ------ Check their altitude answers with your measured known values in far right column of your Table II.

Data Table III (share with other "blind" student group) Ave Vel. =

Data	a Table III	(share with other "blind	" student group) Ave Vel. =	
*R Interva l	Time Average of 3	Original Distance Ball Traveled (Baseline From Data Table I) {D1}	Distance Ball Traveled = (velocity*average time) (cm) {D2}	Altitude (cm) {D1- D2=Altitude}
0 cm	Trials			
20 cm				
40 cm				
60 cm				
80 cm				
100 cm				
120 cm				
140 cm				
160 cm				
180 cm				
200 cm				

### Step 3: Optional Plotting and Graphing the Data (if time, check resolution)

#### **Elaborate**

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### **Follow-Up Questions:**

- 1. Why is it important to keep the distance between each altimeter measurement consistent?
- 2. How could we make the topographical profile more accurate?
- 3. What does the graph look like in the comparison to your model (i.e. the same, inverted etc.)?
- 4. Which looks more like the model, the graph you generated from the shorter or longer distance between readings (intervals)? Why is this?

#### **Evaluate**

### Step 4: Expand your thought process

The Laser Rangefinder aboard NEAR sends out a laser beam and "catches" it as it returns from being reflected by the surface of 433 Eros. The instrument records how long it takes the beam to reach the surface and bounce back. The scientists know how *fast* the beam is traveling; therefore, they can calculate how *far* it traveled. By measuring this time and multiplying by the velocity of the beam, they calculate how far the laser has traveled. They must then divide the distance the beam traveled in half.

1. Ask the students why they did not divide in half to find the distance to the object in *their* topography model.

Next, the scientist must compare this distance to a "baseline" distance we will call zero. On Earth, we might use sea level as the baseline. Another way to set the baselines is to start at the center of the planetary body being studied and draw a perfect circle as close to the surface of the body as possible. Using this baseline, the altitude compared to zero can be calculated and graphed. (Here on Earth, we often say that some point is a certain number of feet above or below sea level).

- 1. Why do we not use the term "sea level" for Mars and other planets?
- 2. The students will now calculate the altitude of the points along their model. To do this subtract the distance the ball traveled at each interval (from Data Table II) from the distance the ball traveled in Step 1 (column B, Data Table I). The number students come up with will be zero or greater. Use Data Table III to do the calculations. The number in column B in this table should be the same for every interval. Remember, it is the baseline altitude and does not change.

\*This exercise was adapted from Goddard Space Flight Center: <a href="http://mola.gsfc.nasa.gov/pingpong.html">http://mola.gsfc.nasa.gov/pingpong.html</a>



#### In-Class Activity 2

Remote Sensing\_REVISED (\*Note: This is a shorter version of the previous activity. Instructors may choose either version depending on time constraints.)

MOLA Simulation\*

Since we ran out of time, we're simplifying the ping pong exercise to just 3 "topographies"- 1) a base datum (floor), 2) a medium level, and 3) the highest level.

For each topography level, the students should have had several timings to ensure that they have a consistent value. Just report the averages here. Transfer data in shade boxes to Data Table III for another group to calculate.

Data Table I (D1 Baseline, datum).

		,	<i>J</i> ·
Level	Distance Ball Traveled	Ave. Time (Seconds)	Velocity (distance/time)
Base			

#### **Data Table II**

Ī	Level	Time Average Calculated Distance Ball		Your known measured height of placed	
		(secs)	(cms) Traveled =	block topography (cm)	
			(velocity*average time)	(keep as your "answer")	
	2- med				
	3- high				

	tear here to give Table III to blind student group	
Transfer the dat	ta shown by the shade areas so they can make the calculations.	

### Data Table III (share with other "blind" student group)

Group	
Give the {D1} Ave Vel. =	

*R Interval	Time Average (secs) of "topographies"	Original Distance Ball Traveled (Baseline From Data Table I) {D1}	Distance Ball Traveled = (velocity*average time) (cm) {D2}	Calculated Altitude (cm) {D1-D2=Altitude}
2- med				
3- high				

After students calculate the altitude of the "unknown" topography heights 2 & 3, check their calculated altitude answers with their measured known values the group had actually measured (their far right column of Table II).

If the calculated doesn't match the measured values, explain why the results might be so different:



Part 4

The Laser Rangefinder aboard NEAR sends out a laser beam and "catches" it as it returns from being reflected by the surface of 433 Eros. The instrument records how long it takes the beam to reach the surface and bounce back. The scientists know how *fast* the beam is traveling; therefore, they can calculate how *far* it traveled. By measuring this time and multiplying by the velocity of the beam, they calculate how far the laser has traveled. They must then divide the distance the beam traveled in half.

1. Ask the students why they did not divide in half to find the distance to the object in *their* topography model.

Next, the scientist must compare this distance to a "baseline" distance we will call zero. On Earth, we might use sea level as the baseline. Another way to set the baselines is to start at the center of the planetary body being studied and draw a perfect circle as close to the surface of the body as possible. Using this baseline, the altitude compared to zero can be calculated and graphed. (Here on Earth, we often say that some point is a certain number of feet above or below sea level).

- 1. Why do we not use the term "sea level" for Mars and other planets?
- 2. The students will now calculate the altitude of the points along their model. To do this subtract the distance the ball traveled at each interval (from Data Table II) from the distance the ball traveled in Step 1 (column B, Data Table I). The number students come up with will be zero or greater. Use Data Table III to do the calculations. The number in column B in this table should be the same for every interval. Remember, it is the baseline altitude and does not change.

\*This exercise was adapted from Goddard Space Flight Center: <a href="http://mola.gsfc.nasa.gov/pingpong.html">http://mola.gsfc.nasa.gov/pingpong.html</a>



#### Homework 1

Remote Sensing\_MFE Google Mars-Following Opportunity

**Objective:** The purpose of this homework set is to get you familiar with different types of Mars remote sensing imagery and programs.

Google Mars-Following Opportunity

#### **Directions/Questions:**

Download Google Earth if you haven't already: <a href="http://www.google.com/earth/download/ge/agree.html">http://www.google.com/earth/download/ge/agree.html</a>

In the icon list across the top of the window click on the planet with a single ring and a small dropdown arrow. The students should see options for Sky, Mars, and Moon. Click on Mars.

1.	Name the 5 types of spacecraft imagery available through Google Mars.
	What do the 5 acronyms stand for?
	a.
	b.

c. d.

e.

Find Olympus Mons (see if the students can find it without typing the name in the "fly to" box).

- 2. What is the highest elevation according to Google (find the appropriate *Global Map Layer* in order to determine this information)?
- 3. In the *Global Maps Layer*, besides the *Visible Imagery*, which imagery gives the students the highest resolution of the volcano? Why is this the case?



Go to the Rovers and Landers layer

4. What are the current coordinates of these 3 lander sites? Phoenix Lander Viking 2 Lander Mars 3 Lander

5. Where did MER Opportunity Rover land? (i.e. what crater?)

What crater did it visit next?

Look at the Burns Cliff panorama photo (camera icon, students may have to click on a couple to figure out the right one).

6. List 2 observations the students can make about the photo (colors, shapes, lineations, etc.)?

Name 2 other craters the Opportunity rover explored.

- 7. Write down two observations about what students see in the bottom/ centers of Victoria Crater. Can students name the features?
- 8. Using the Traverse Path layer of the MER Opportunity Rover, locate its position on Sol 1685 (sol= Mars day). What annotated feature (labeled named) is it nearest?



#### Homework 2

Remote Sensing\_MFE *Mars Image Analysis* 

**Directions**: View the following THEMIS image and answer the questions about the image. As students view the image, think about how this image might support one or more of NASA's main exploration goals:

- 1. Determine if life ever existed on Mars
- 2. Characterize the climate of Mars
- 3. Characterize the geology of Mars
- 4. Prepare for future human exploration of Mars

Go to: http://themis.asu.edu/

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*Getting to know THEMIS imagery (click on the "about")* 

- 1. In a few sentences explain what THEMIS detects and how it works.
- 2. Go to THEMIS image: <a href="http://themis.asu.edu/node/5765">http://themis.asu.edu/node/5765</a> What is the title of the THEMIS image?
- 3. Study the THEMIS image. List at least two features you observe.

a.

b.

4. If the sun is illuminating from the left, are the features expressing positive (hill) or negative (valley) relief? If features differ from another (i.e. one has positive relief and the other negative) describe their relief separately.

5.	What is the Lat/Long of the ce	nter of THEMIS image?
	Lat	Long



- 6. Explain how this image meets or does not meet NASA's exploration goals of Mars.
- 7. If students were to lead a lander mission to an area located within the image, where would they land and why?

### More THEMIS Imagery

- 8. Go to the THEMIS image gallery by Topic: <a href="http://themis.asu.edu/gallery">http://themis.asu.edu/gallery</a>
  Have the students choose an image they like and report the following:
  - a. What is the image ID or the image url?
  - b. Why did the students choose this image?
  - c. Where is the image located?
  - d. Near what major Mars geographic region is it located (South/North pole, Victoria Crater, Endurance Crater, Merdiani Planum, Hellas Basin, etc.)? Use the *View this image on Map* link at the bottom of the image data column to see a map view of Mars.
  - e. Why might this location be important to science?



### HiRISE Imagery

- 9. Navigate to the HiRISE website: <a href="http://hirise.lpl.arizona.edu/">http://hirise.lpl.arizona.edu/</a>
  - a. Scroll to the bottom of the page (gray box) and click on the link "Science Themes". Click on the *Aeolian Processes* file of images. Under the main image click "View Images in this Theme." Find image titled "Dunes in the Western Nereidum Montes." If students cannot find the image type ESP\_013046\_1390 into the search box.
  - b. Define the term *Aeolian*. (also known as eolian)
  - c. Why might an image of *aeolian* processes on Mars be of interest to us on Earth?
  - d. Have students sketch what they see below. Label appropriate parts (high and low areas). Can they identify the direction of the wind if North on Mars is up? If so, what direction (cardinal direction) is it?

